



## The University of Western Australia, School of Physics Frequency and Quantum Metrology Research Group

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- ↘ [Eugene Ivanov](#)
- ↘ [John McFerran](#)
- ↘ [Sascha Schediwy](#)
- ↘ [Jean-Michel Le Floch](#)
- ↘ [Yaohui Fan](#)
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- ↘ [Stephen Parker](#)
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- ↘ Romain Bara-Maillet
- ↘ Jeremy Bourhill
- ↘ Nikita Kostylev
- ↘ Natalia Carvalho
- ↘ Akhter Hoissan





## TESTING LORENTZ INVARIANCE AND FUNDAMENTAL CONSTANTS WITH PRECISION CLOCKS AND OSCILLATORS

**ME Tobar<sup>1</sup>, SR Parker<sup>1</sup>, PL Stanwix<sup>1</sup>, JJ McFerran<sup>1</sup>, EN Ivanov<sup>1</sup>, J Guéna<sup>2</sup>, M Abgrall<sup>2</sup>, S Bize<sup>2</sup>, A Clairon<sup>2</sup>, Ph Laurent<sup>2</sup>, P Rosenbusch<sup>2</sup>, D Rovera<sup>2</sup>, G Santarelli<sup>2</sup>, M Nagel<sup>3</sup>, E Kovalchuk<sup>3</sup>, A Peters<sup>3</sup>**

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- 1) Future plans in Western Australia for contribution to ACES
- 2) Rotating Michelson Morley Experiment from UWA to Berlin
- 3) Experiments at SYRTE Paris



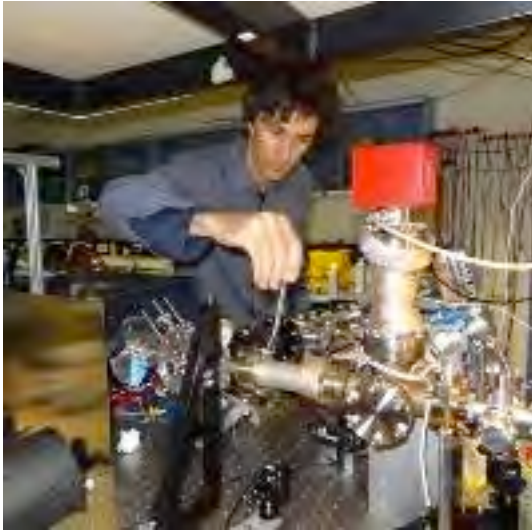
## UWA Contribution to ACES: Past Contributions

- ↘ Since 2000 the Australian Research Council has supported the FSM labs on research related to ACES
- ↘ Developed precision frequency technology
- ↘ Developed Expertise in testing Fundamental Physics
- ↘ 2000-2012 \$4.5M for 8 projects in related research
  
- ↘ LONG TERM GOAL : To become the Southern Hemisphere hub for Space Missions involving precision clocks
  
- ↘ Complicated by no space Agency, but strong support in Australia through local funding agencies



## UWA Contribution to ACES: Current Contribution

- **Four ARC Projects from 2009 – 15 worth \$5.1M**
- **UWA to Further Upgrade Labs in 2013 by \$0.5M (Spent \$0.8M)**
  
- ARC Linkage Infrastructure LE110100054 2011-13 Funding ARC \$1,230,000  
Universities \$601,917
- To buy/develop infrastructure for ground station: JPL clocks and Yb Lattice clock
- Tobar ARC Laureate Fellowship FL0992016 2009-1 Funding \$2,200,000 ARC  
Salaries for Tobar, Postdoc, Technician plus project money
- Future Fellowship (John McFerran) FT110100392 2012-2015 Funding  
\$540,000  
Salary for McFerran plus project money
- ARC Discovery DP130100205 2013-15  
Funding \$500,000  
Salary for Postdoc and project money to use ACES ground clocks at UWA to test  
fundamental physics



## Ytterbium Lattice Clock

<http://www.news.uwa.edu.au/201310046103/research/ultimate-accuracy-machine>

- Vacuum system is fully assembled including the Yb oven (effusion cell), Zeeman slower, main chamber and ion + getter pumps.
- The Zeeman slower is in operation (with 399nm): the most probable velocity for the slowed group of atoms is less than 10m/s (significantly less than the designed captured velocity of the MOT ~ 50m/s)
- Can see slowing on all the bosonic and relevant fermionic isotopes
- The MOT coils are in place and maintain current.
- Presently setting up the laser beams for the MOT (399nm).
- Conservative estimate by the end 2015 -> laser locked to the clock transition (with lattice trapping), possibly sooner.



# Currently at UWA site Prior funding from ARC

GPS antenna



2.4 m antenna TWSTT

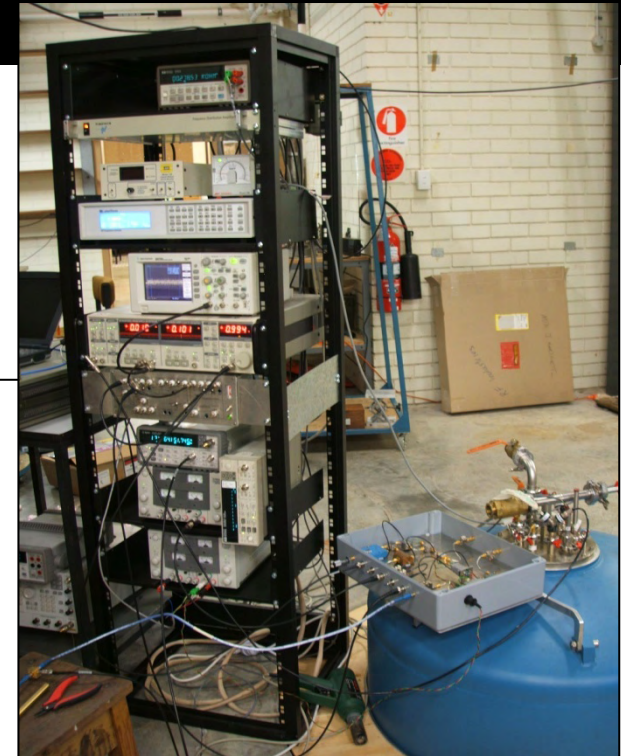


**Kvarz H-Maser**

$f_{out} = 100 \text{ MHz}$   
 $\sigma_y(\tau) \sim 2 \cdot 10^{-15}$   
 $\tau > 10^3 \text{ s}$   
drift  $\sim 3 \cdot 10^{-16} / \text{day}$



Synthesis  
between  
H-Maser and  
CSO



**CSO**

$f_{out} = 11.200 \text{ GHz}$   
 $\sigma_y(\tau) < 1 \cdot 10^{-15}$   
 $\tau < 1000 \text{ s}$   
drift  $\sim 2 \cdot 10^{-15} / \text{day}$

New funding -> ACES Ground Station and Atomic Clocks



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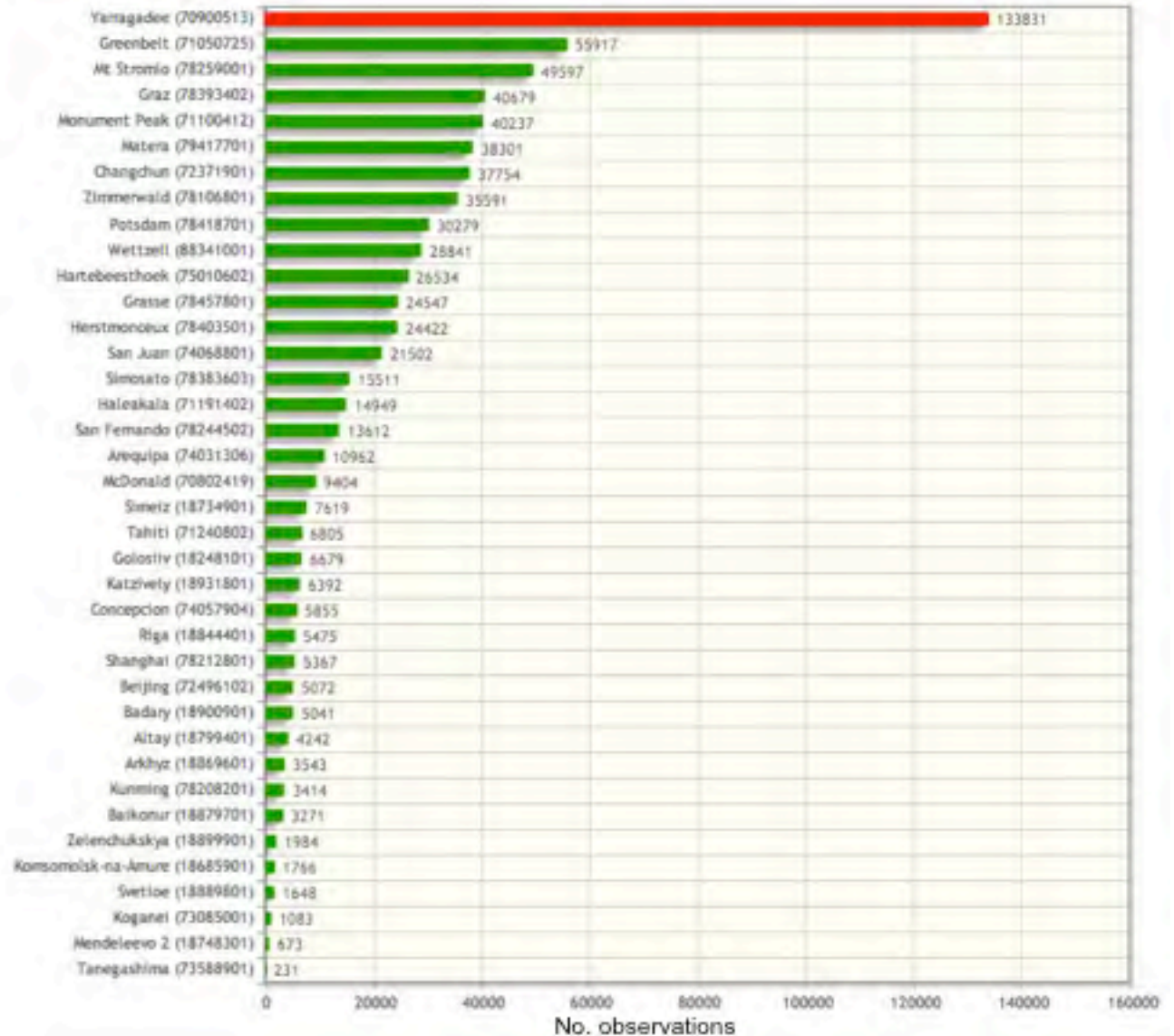
# Reorganize Roof to install the Ground Station



# Yarragadee Laser Ranging Station



Stations vs. Observations





# Transfer over optical fibre





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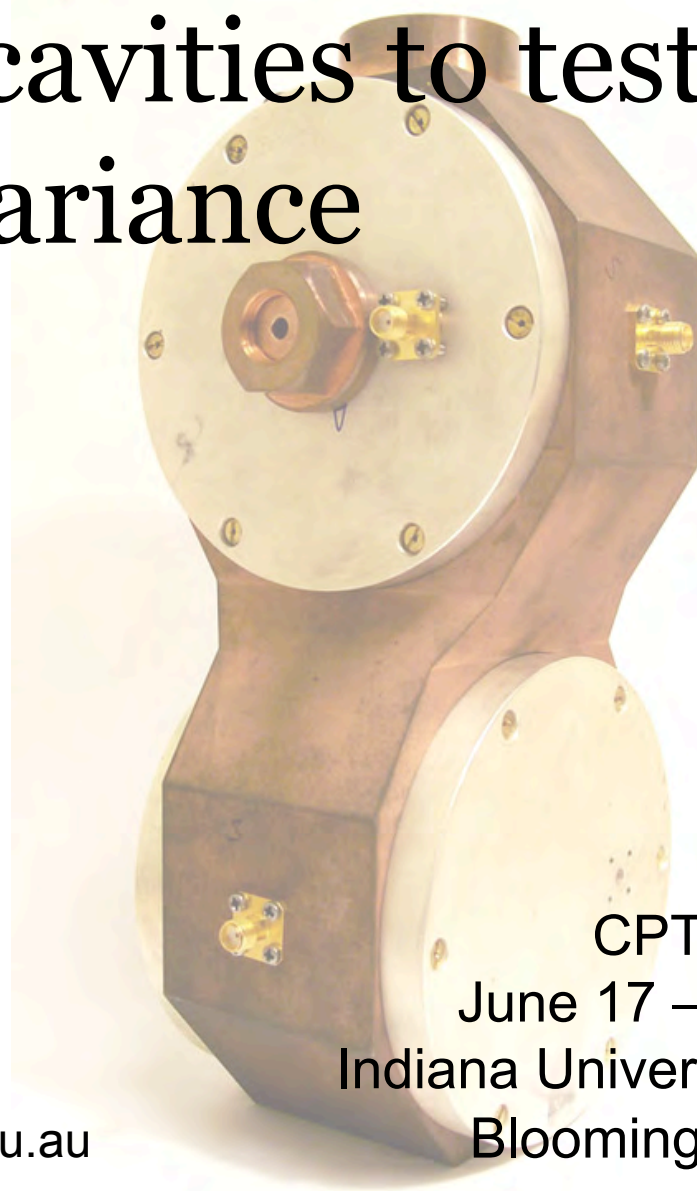
# Using microwave cavities to test Lorentz invariance

*Stephen Parker, Paul Stanwix, Eugene  
Ivanov and others*

Frequency & Quantum Metrology  
School of Physics, UWA

*Moritz Nagel, Evgeny Kovalchuk,  
Achim Peters and others*

Quantum Optics & Metrology  
Institute for Physics, HUB



CPT 13

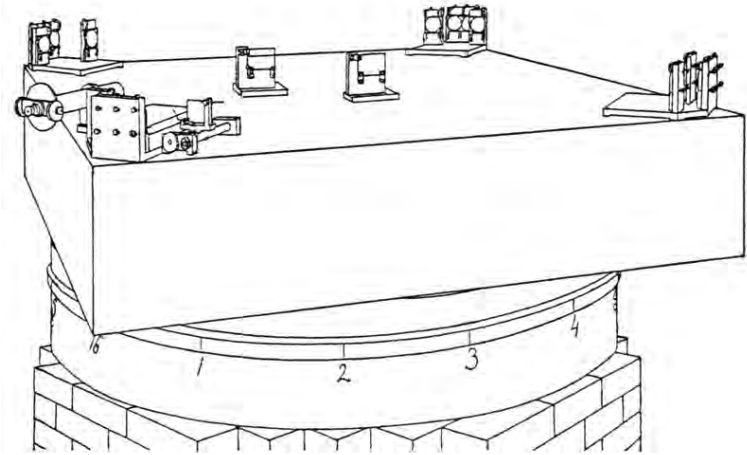
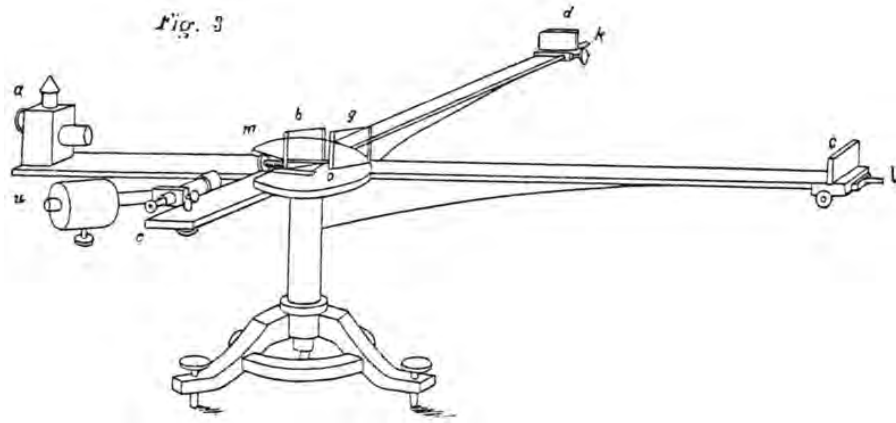
June 17 – 21

Indiana University

Bloomington



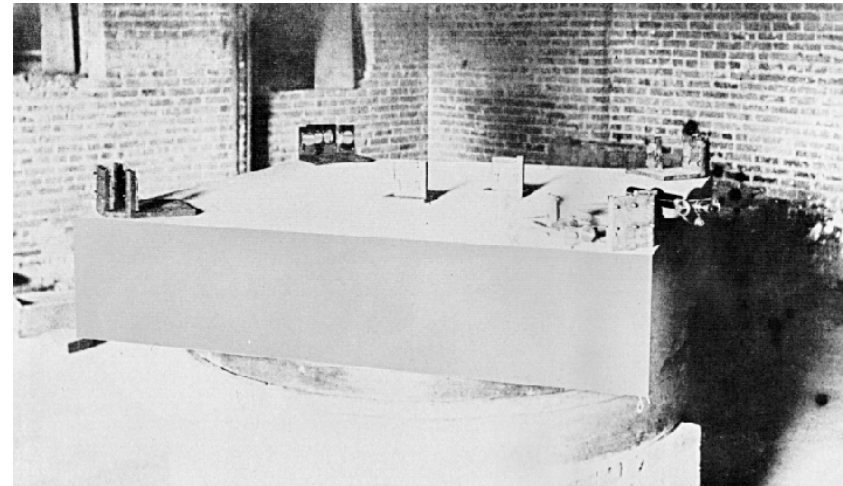
[Stephen.Parker@uwa.edu.au](mailto:Stephen.Parker@uwa.edu.au)



Light goes this way

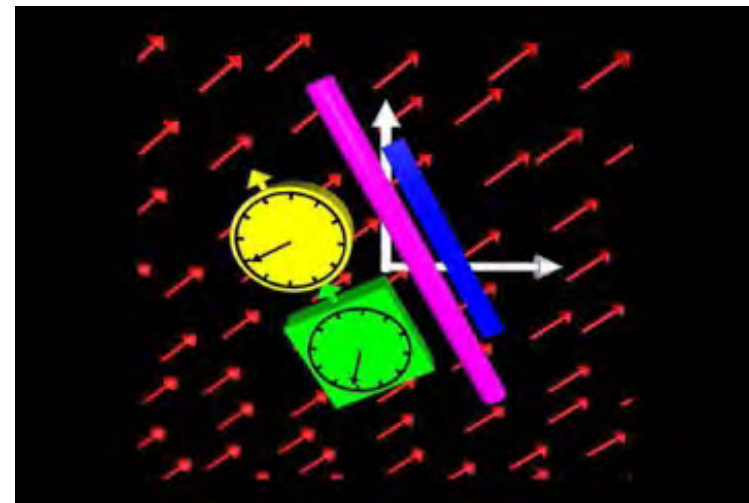
Light goes that way →

Which way was faster?



# Testing Relativity: Standard Model Extension (SME)

Ansatz: Extend the Standard Model of elementary particle physics by all Lorentz violating terms which leave observer transformations invariant.



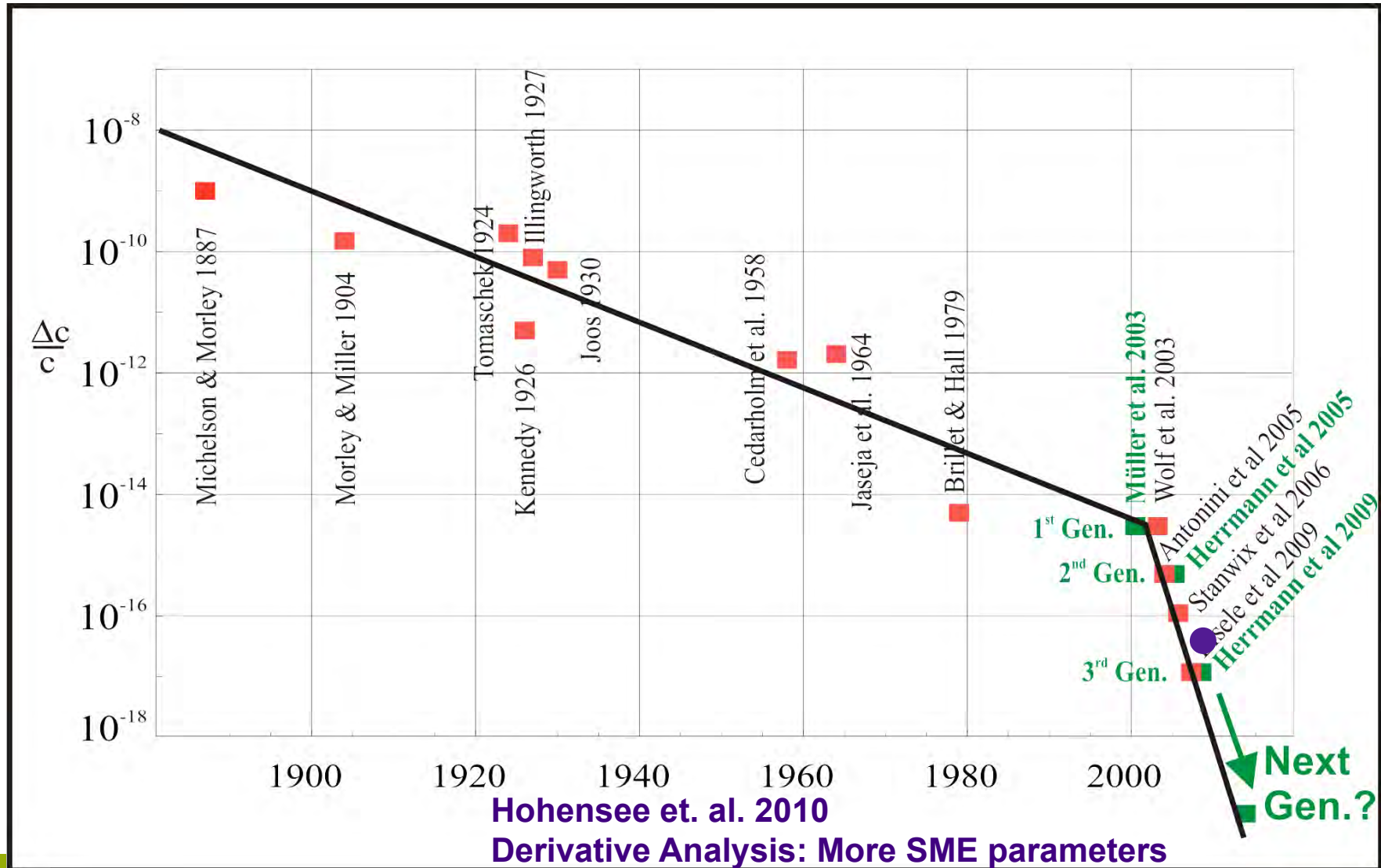
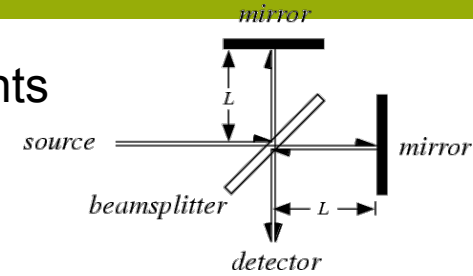
Example  $\mathcal{L} = \underbrace{-\frac{1}{4}F^{\mu\nu}F_{\mu\nu}}_{\text{Standard}} - \underbrace{\frac{1}{4}(k_F)_{\kappa\lambda\mu\nu}F^{\kappa\lambda}F^{\mu\nu}}_{\text{Lorentz Invariance}}$

The magnitude or vanishing of these terms should give hints for structure of an unified theory!

Colladay & Kostelecký, Phys. Rev. D55, 6760 (1997),  
 Colladay & Kostelecký, Phys. Rev. D58, 116002 (1998)

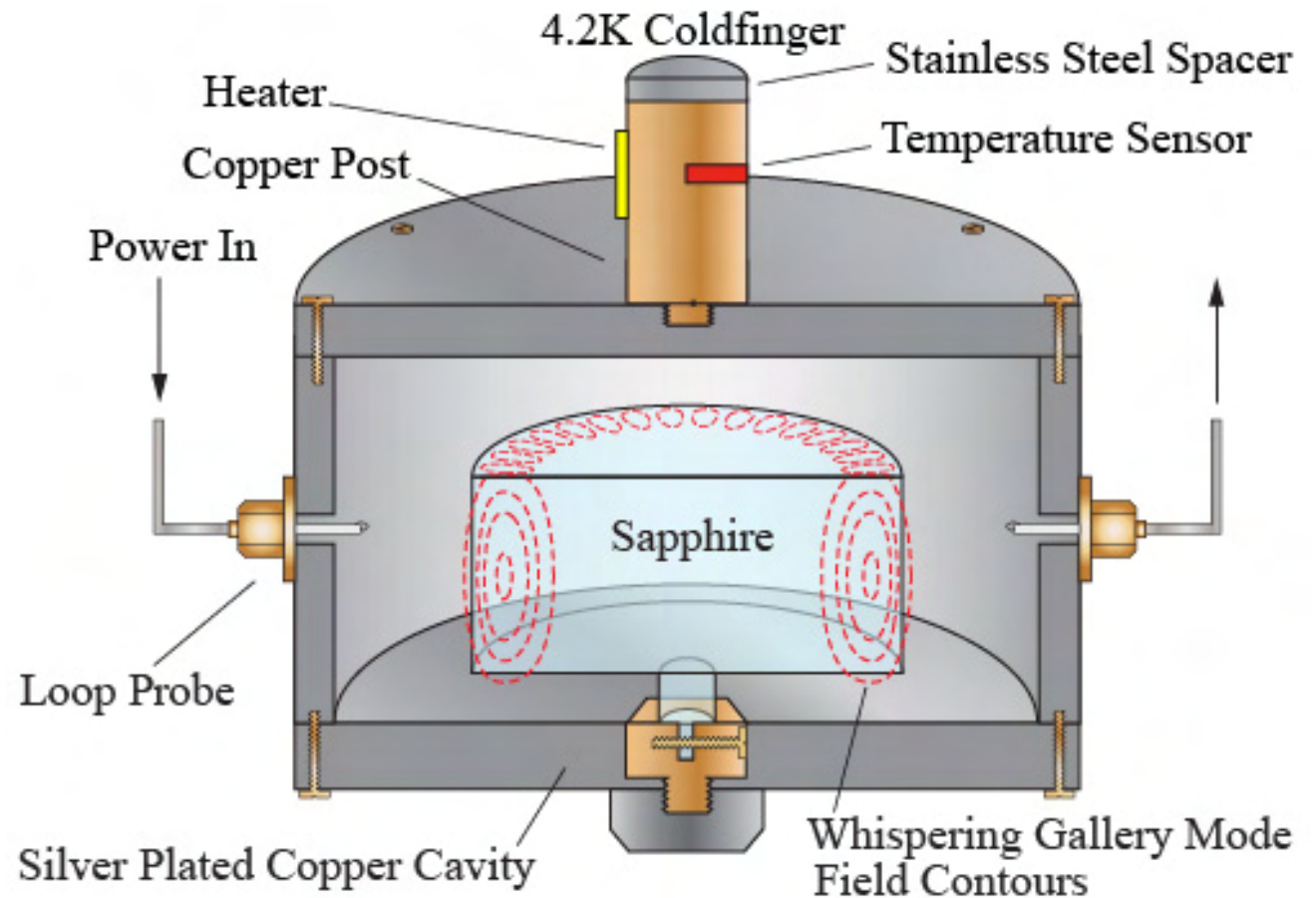
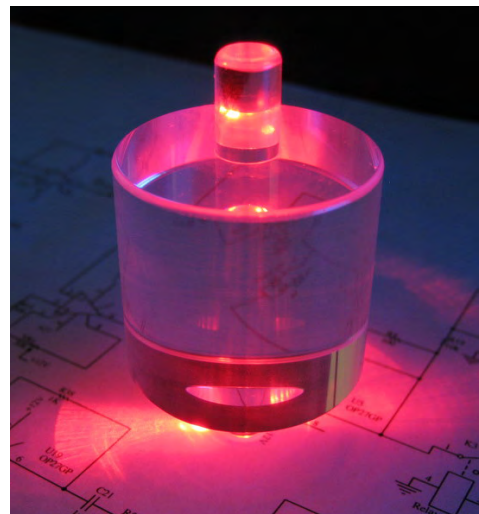
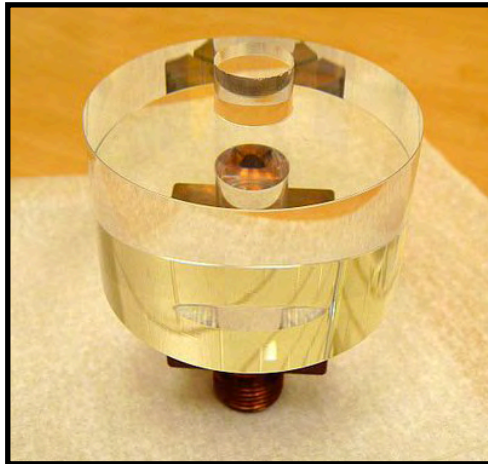
# Rapid Improvements in Michelson-Morley Experiments

Next version will probe the Plank Scale



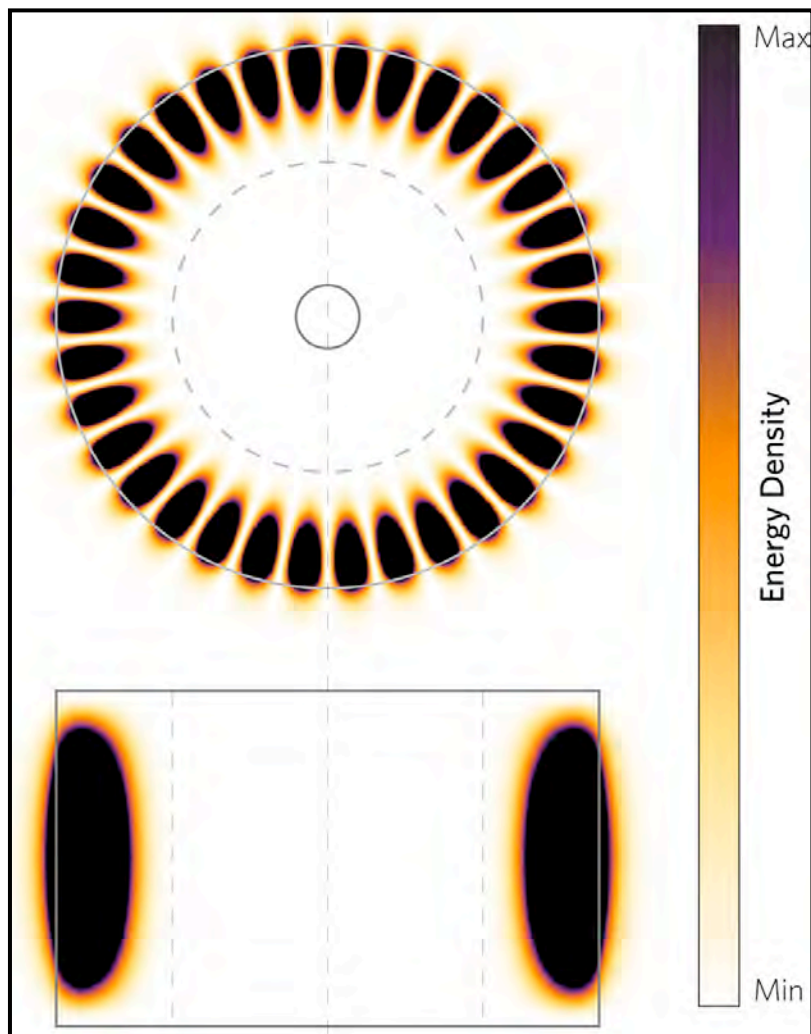


## Sapphire Loaded Cavity Resonator



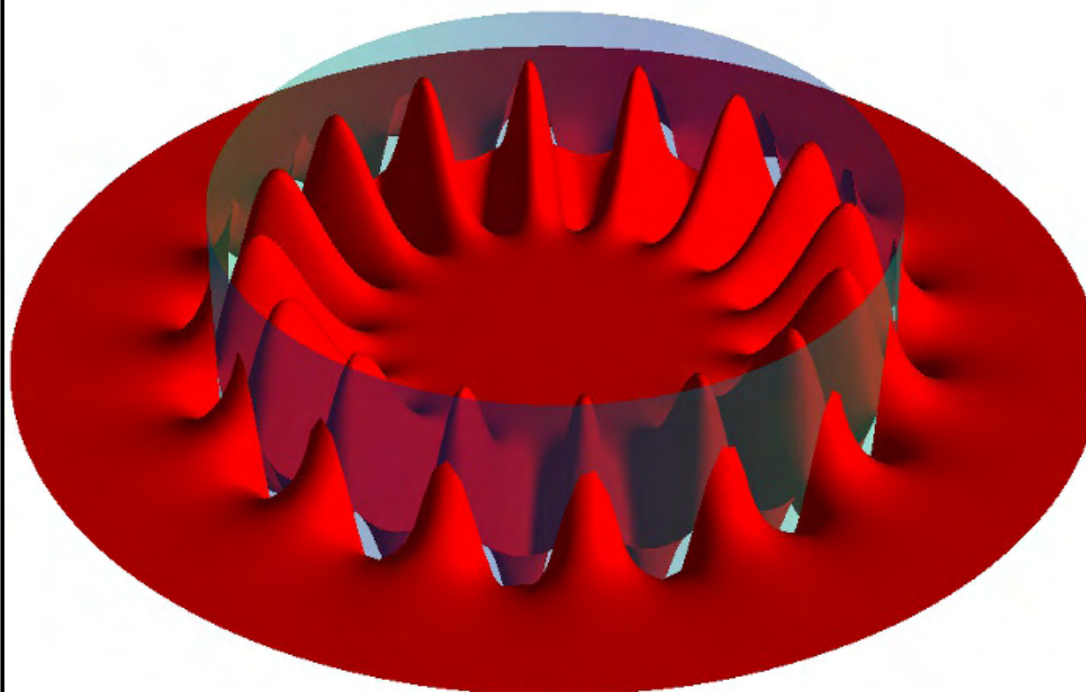


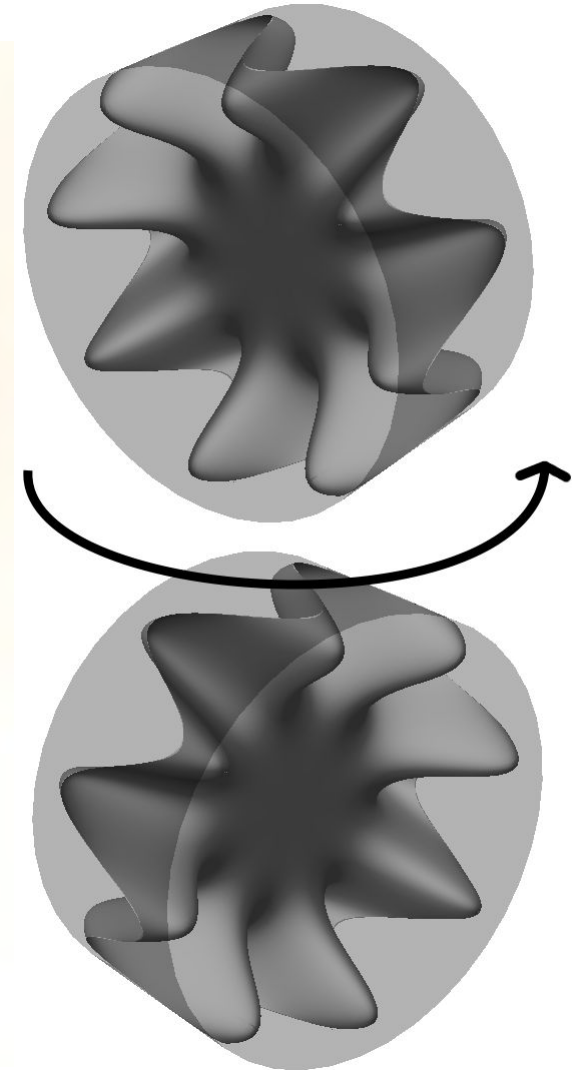
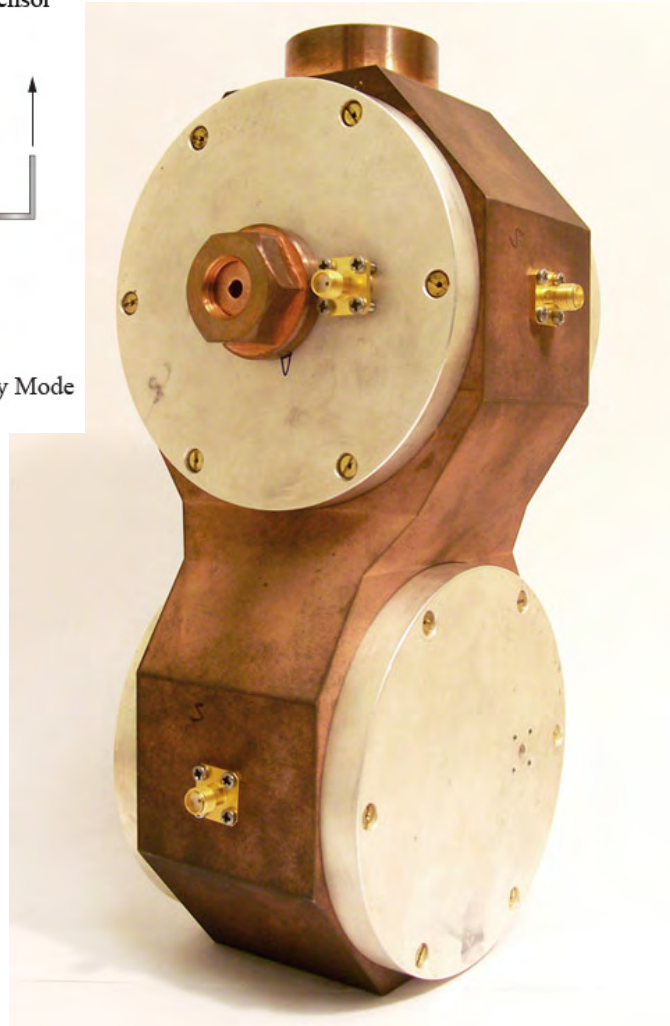
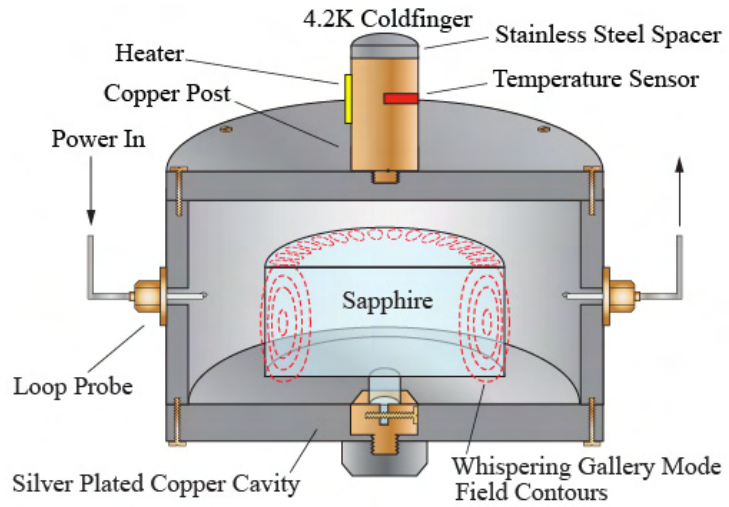
## Whispering Gallery Modes in Sapphire



$WGH_{x,1,1}$  – transverse magnetic,  $E_z$  and  $H_r$   
 $WGE_{x,1,1}$  – transverse electric,  $H_z$  and  $E_r$

WGE is more sensitive due to larger magnetic filling factor

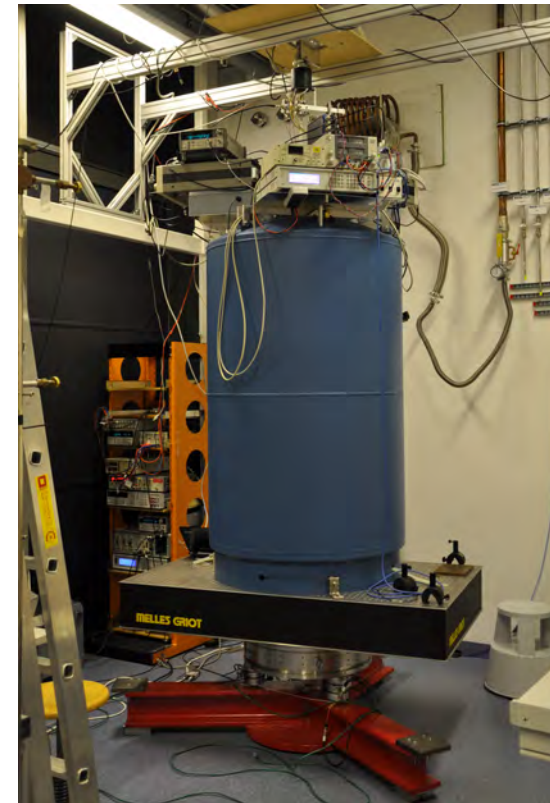
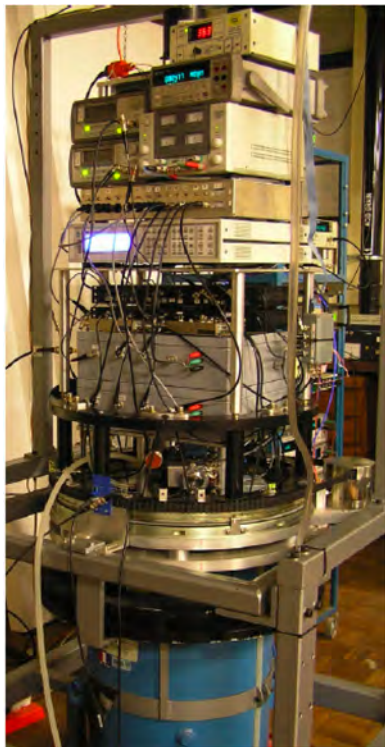


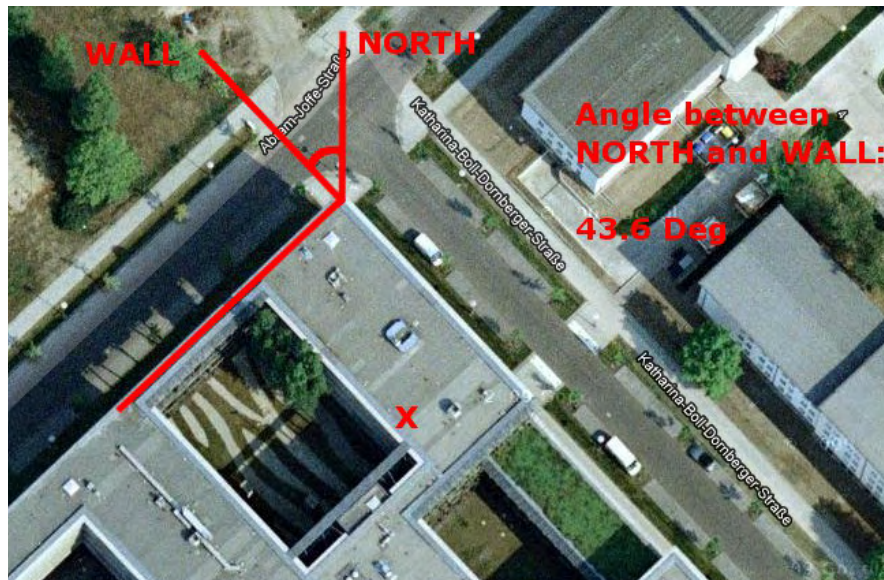
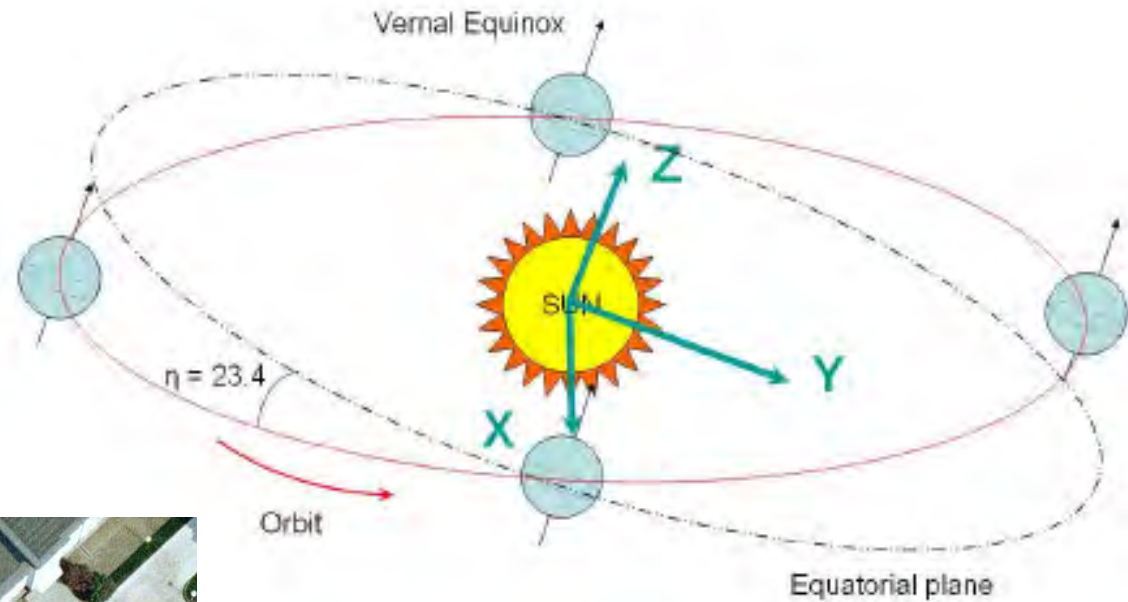






## UWA rotating Michelson-Morley experiment re-visited -> Berlin





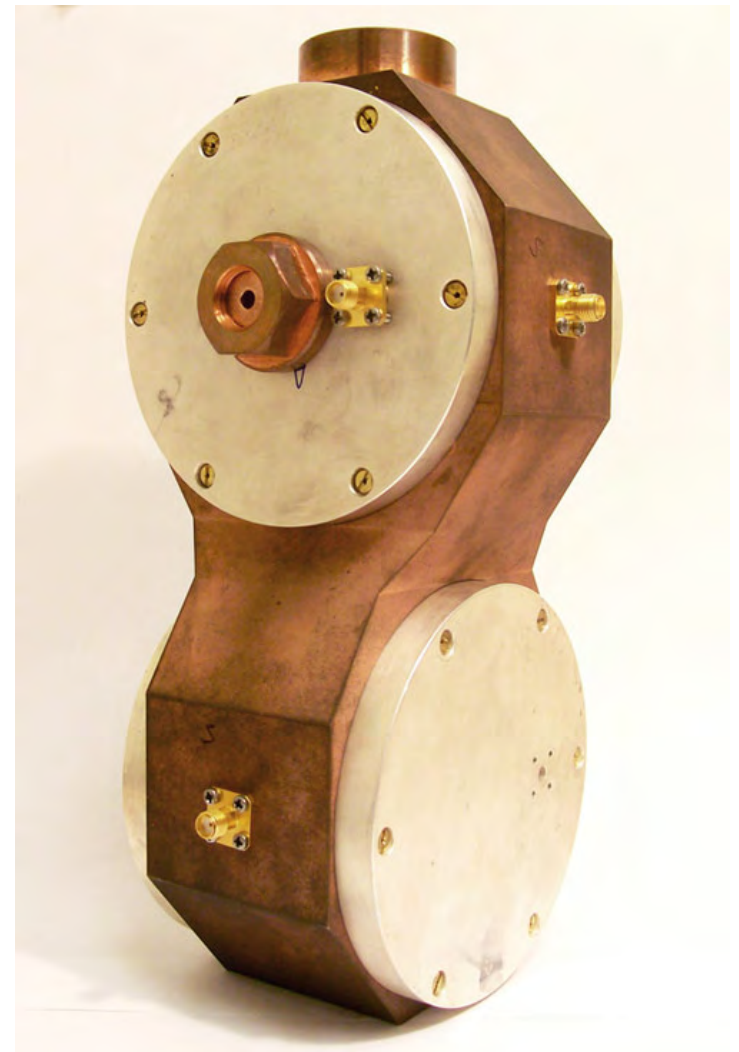
# New Experiment -> Berlin

New dual cavity design (right) allows for better thermal stability

Larger sapphire crystals have a higher quality factor  $\approx 2 \times 10^9$  (compared to  $2 \times 10^8$  the prior experiment) allows better frequency stability

$WGE_{16,0,0}$  mode is more sensitive to Lorentz violating parameters ( $S = 0.4567$  compared to  $S = 0.1958$ ) -> over all near two orders of magnitude improvement.

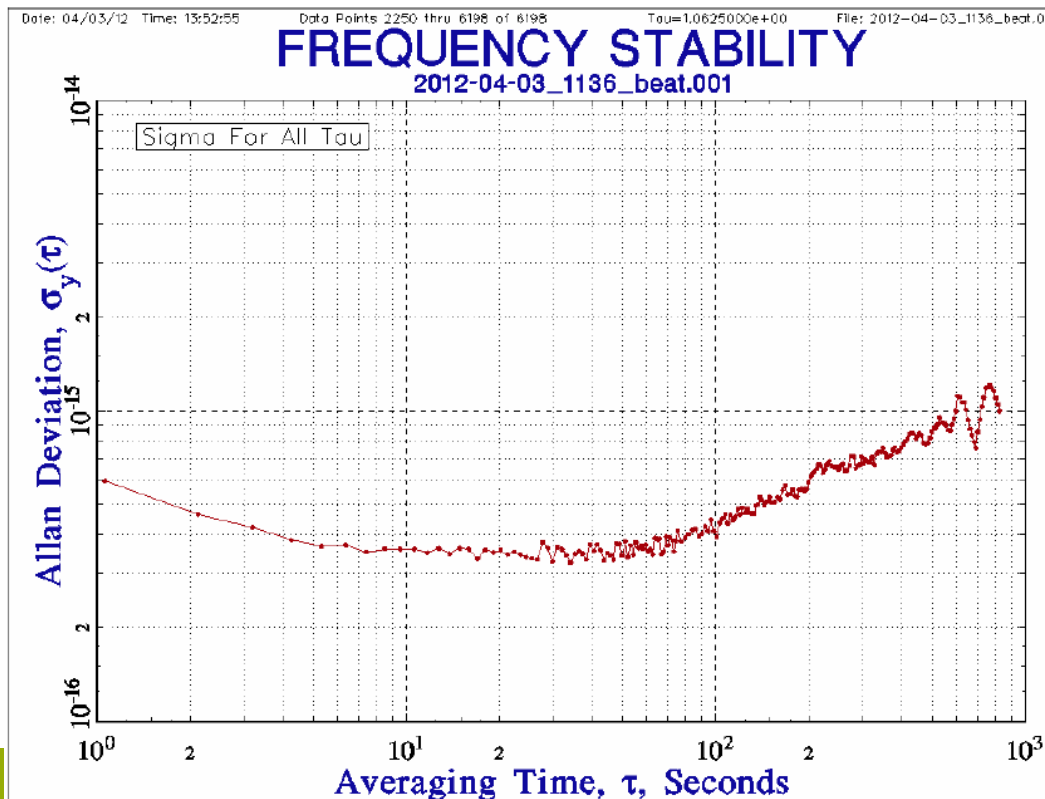
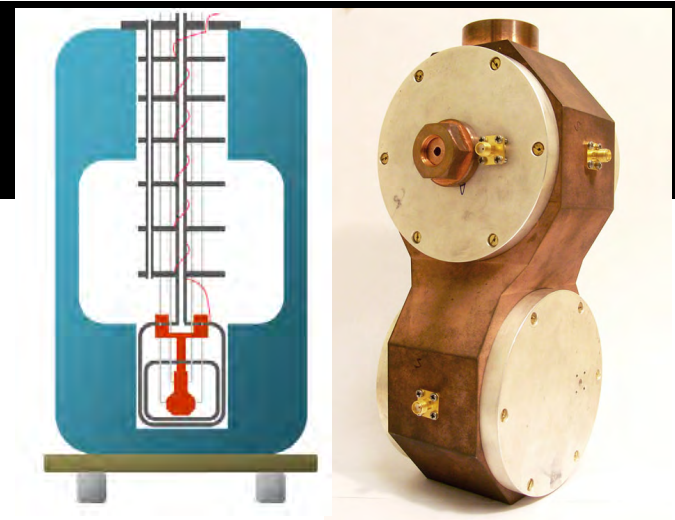
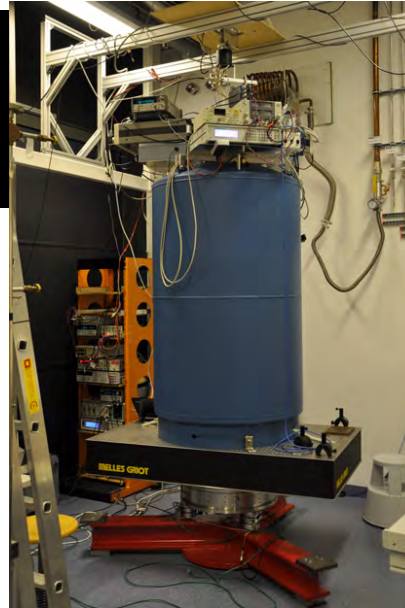
Reduction of noise-inducing systematics (i.e. tilt) use Berlin system.





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Cryogenic Sapphire Oscillators  
in Berlin, started operation  
August 2011  
Ready for new limits  $\sim 10^{-18}$



Long process of  
eliminating hard to  
diagnose systematic  
errors and bugs in  
data acquisition.  
Have 1.25 years data ->  
 $\sim 10^{-18}$  in the photon  
sector



# Testing local position and fundamental constant invariance due to periodic gravitational and boost using long-term comparison of the SYRTE atomic fountains and H-masers

M. E. Tobar,\* P. L. Stanwix, and J. J. McFerran

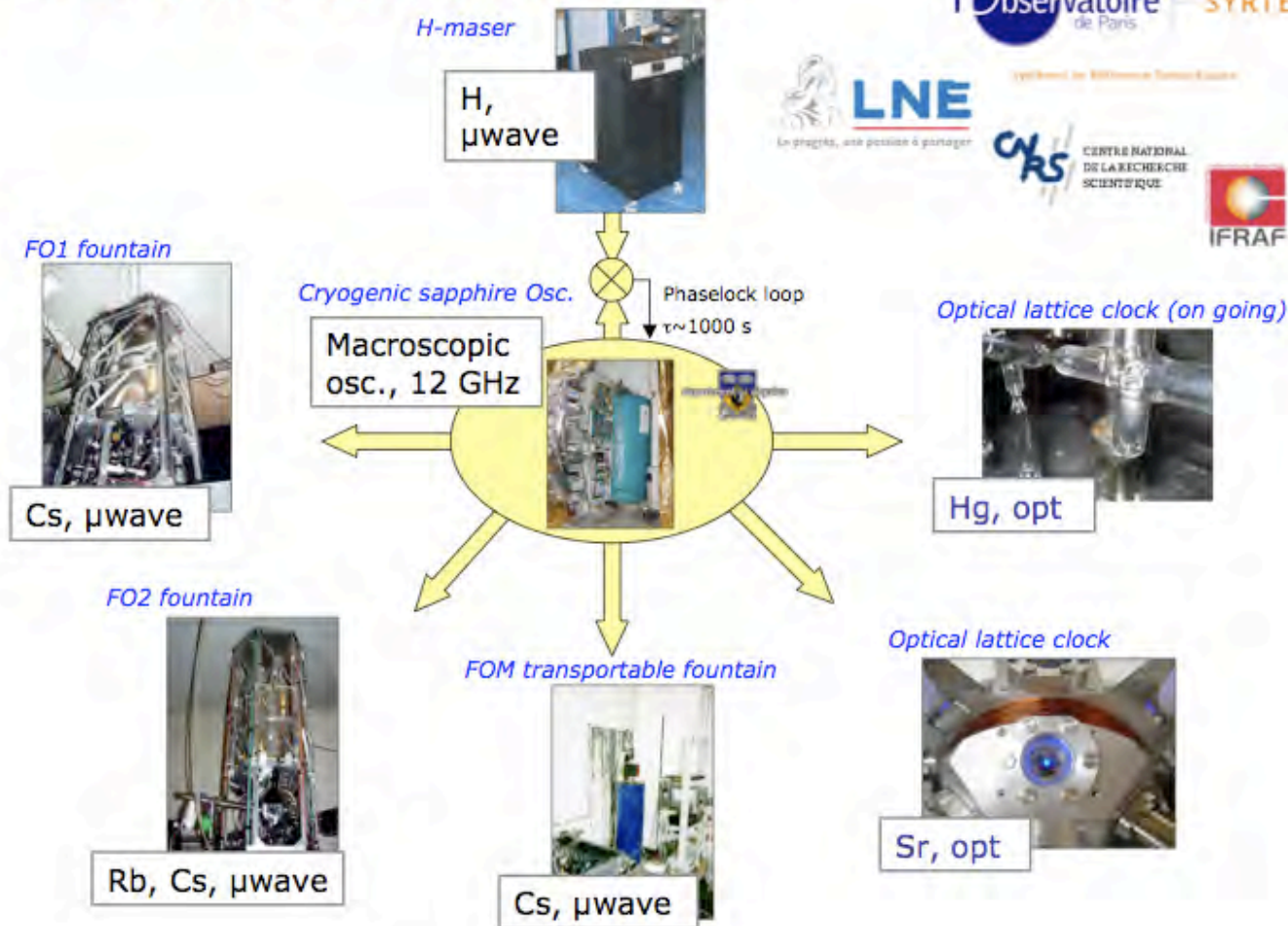
School of Physics, University of Western Australia, 6009 Crawley, Australia

J. Guéna, M. Abgrall, S. Bize, A. Clairon, Ph. Laurent, P. Rosenbusch, D. Rovera, and G. Santarelli

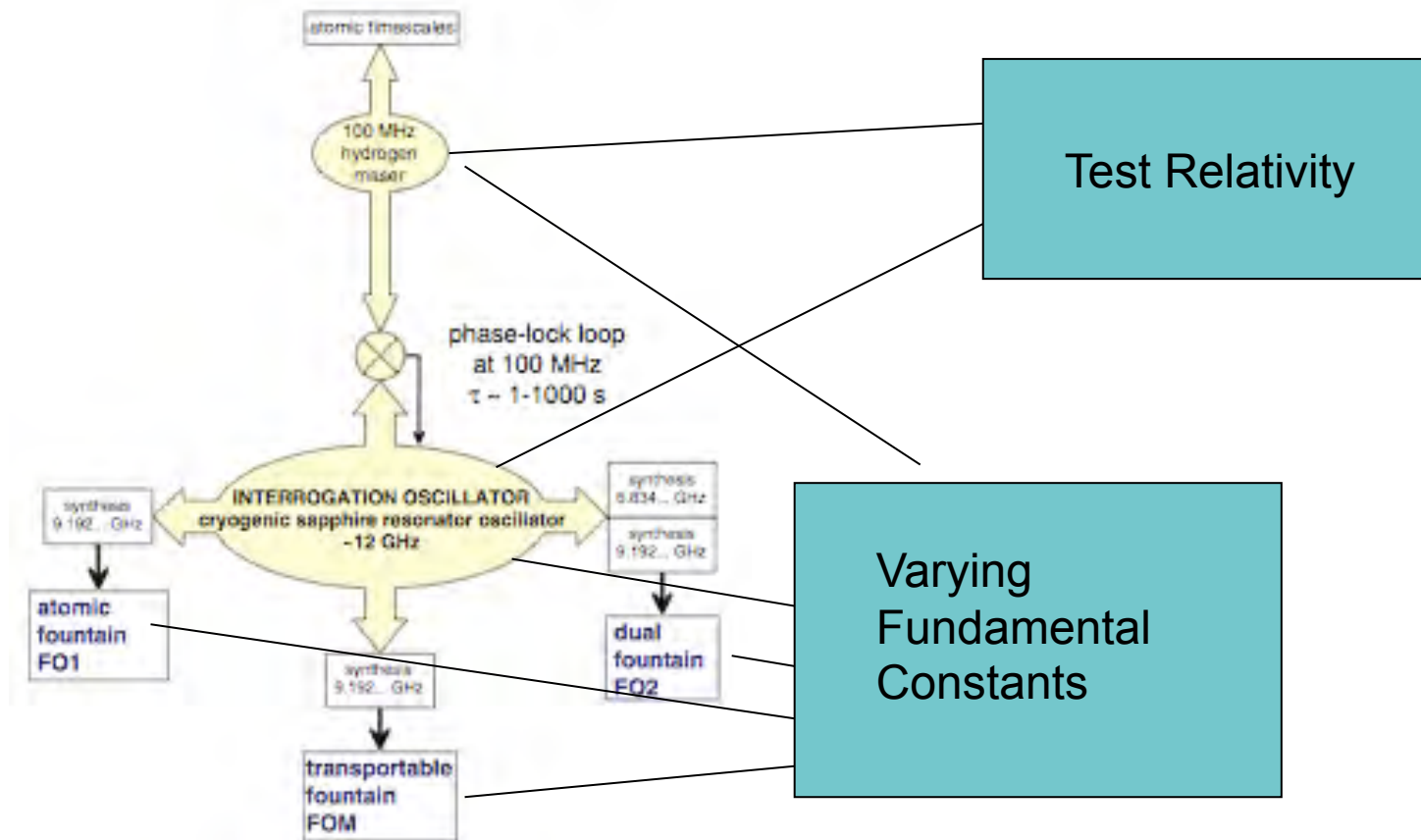
LNE-SYRTE, Observatoire de Paris, CNRS, UPMC, 75014 Paris, France

(Received 28 March 2013; published 5 June 2013)

## LNE-SYRTE CLOCK ENSEMBLE



$$x_{1-2} = \frac{\nu_1}{\nu_2} = \text{const} \times \alpha^{n_\alpha} \mu_e^{n_e} \mu_q^{n_q}$$



This work: Test 3 Cs and 1 Rb Fountains against H-maser

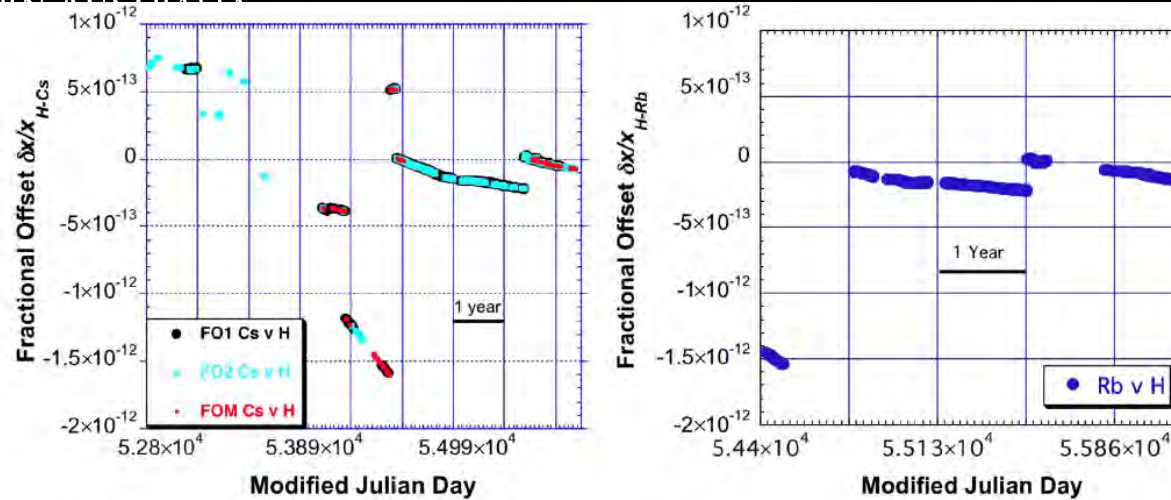


FIG. 1 (color online). Left: Measured fractional offset variations of the frequency ratio between three Cs fountains (FO1, FO2, and FOM as identified by the legend) and various H-masers. Long-term results span from 2/7/2003 to 2/11/2011, which is 3,045 days (8 yr and 4 months). Right: Measurement between the FO2 Rb fountain and various H-masers over nearly a 5 yr span.

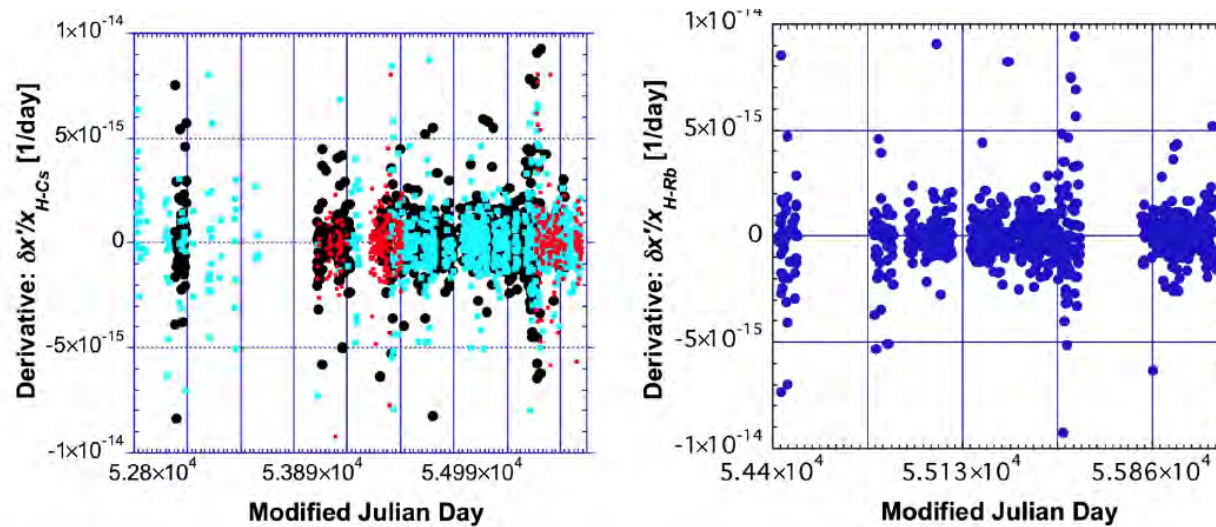


FIG. 2 (color online). Left: The derivative with respect to time of the Cs vs H data shown in Fig. 1, in units of fractional frequency per day. Each frequency measurement is averaged over 95,000 sec (1.1 days) before the derivative is taken. Right: Rb vs H data after following the same procedures. This data are used for searching for variations at the annual period.



“null red-shift experiment”  
Due to change in sun potential

“null Doppler-shift experiment”  
With respect to CMB  
“Cosmic Doppler Shift”

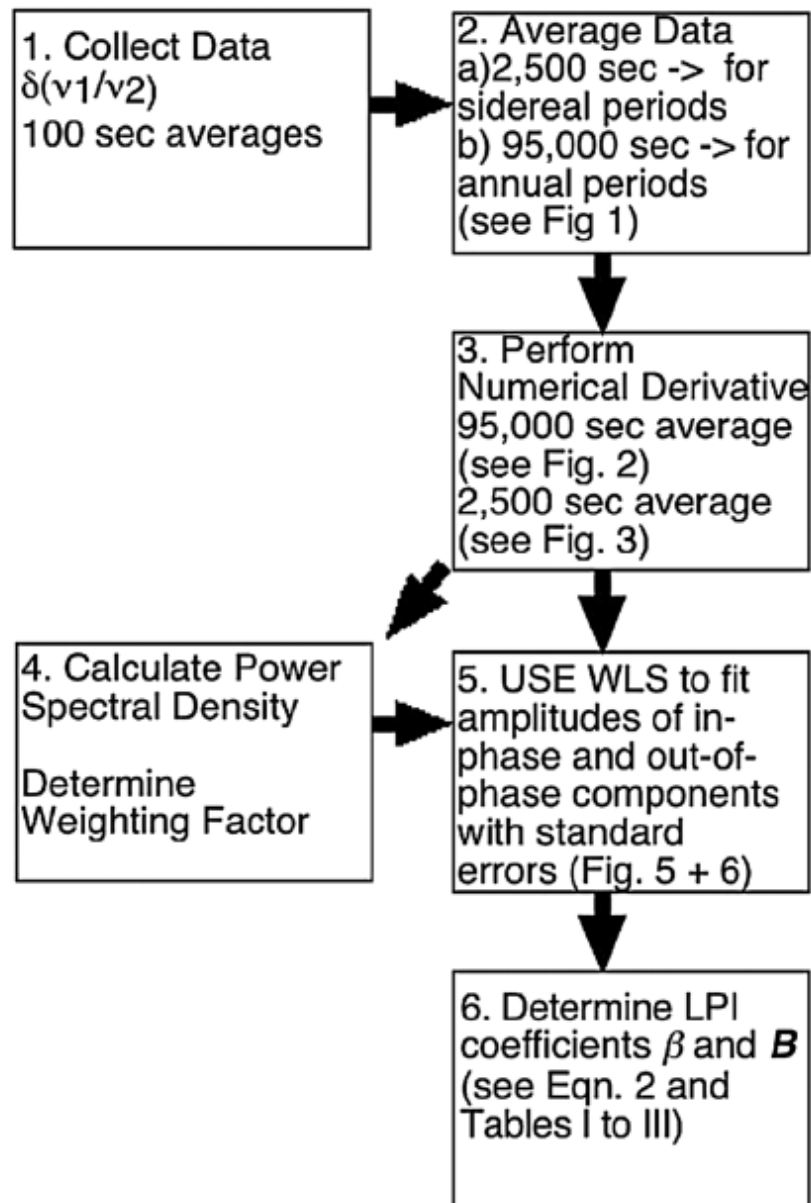






TABLE II. The data in Table I are overconstrained by the measurements. Thus, a least squares analysis was implemented to determine each component of the boost-violation vector with the associated standard error.

Boost-Violation vector component	H-maser/CSO [24]	Cs/H-maser	Rb/H-maser
$B_x$	$-12.6(9.1) \times 10^{-11}$	$1.6(1.1) \times 10^{-11}$	$-2.2(2.6) \times 10^{-11}$
$B_y$	$6.5(9.8) \times 10^{-11}$	$-7.6(3.3) \times 10^{-11}$	$-0.8(3.6) \times 10^{-11}$
$B_z$	$-46.8(48.5) \times 10^{-11}$	$-15(15) \times 10^{-11}$	$-5.6(9.7) \times 10^{-11}$

TABLE III. Decomposition of the boost-violation vectors in Table II to limits on the invariance of fundamental constants with respect to boost. The values range from  $10^{-10}$  to a few parts in  $10^{-9}$ .

Fundamental constant boost-violation vector	$i = x$	$i = y$	$i = z$
$B_{\alpha i}$	$1.1(0.9) \times 10^{-10}$	$-1.6(1.4) \times 10^{-10}$	$-2.0(4.8) \times 10^{-10}$
$B_{ei}$	$-5.3(3.4) \times 10^{-10}$	$6.2(5.2) \times 10^{-10}$	$1.4(18) \times 10^{-10}$
$B_{qi}$	$-7.3(6.7) \times 10^{-10}$	$5.9(9.8) \times 10^{-10}$	$1.4(29) \times 10^{-10}$

# THE END

A digital chalkboard with a dark blue background and an orange border. It features several mathematical formulas:  $\Delta = 0.07 + \pi \sqrt{m}$ ,  $\sqrt{xy} = z$ ,  $e = \pi \sqrt{.03}$ ,  $\sqrt{xy} = z$ ,  $E = MC^2$ ,  $\pi = .03 \sqrt{\frac{14}{\pi}}$ ,  $\sqrt{xy} = \Delta$ , and  $= -b$ .

